Physics 303 Classical Mechanics I Aradudian & Review William & Mary A.W. Jachura -

## Introduction & Review

Classical Mechanics (CM) is concorred with the Indy I mation I bodies. It serves is the backbare I modern playsics, relating simple physical laws to complicated dynamics exhibited by mechanical systems. This course builds on the foundations you learned in Phys. 208. We will apply Newton's laws l'hyragia rechaires to systems I particles, rigid bodies, & continuous media. De will examine the behavior I systems in non-mention reference frances, & explore le cursequences f chartic motion from nonlinear systems. Advanced mathematical techniques such as perturbution theory to approxinde the dynamics of convicted systems.

Finally, we will introduce another formulation of mechanics, Mamiltonian Mechanics. Harritonian Mechanics provides a system Tix framework to examine the phase space of dynamic unides, which impacts how are formulates Quatum Uncorics.

Newtarian Mechanics

CM is built upon Neiton's laws & motion. Hore we first consider a structurelass body (a particle) Which exists in 3D space (ess; (x,7,2)) & time (t). In CM, there is a universal time for which all observes agree upon.

NI I some incidial refrance trune  
such at it there are no external forces,  
$$\vec{F} = \vec{O}$$
, ading a a baby, then  
 $\vec{V} = \vec{O}$   
 $\Delta_{ij} = \vec{O}$ 

Hoctore the object moves it a right line,  

$$\vec{r}$$
 reforme the  
 $\vec{r}$  it) =  $\vec{r}$ , +  $\vec{v}$  (t-t.)  
position with  $\vec{r}$  it) position,  $\vec{r}$  (to) =  $\vec{r}$ .  
some chosen  
 $\vec{v}$  =  $\vec{r}$ 

We specify a institut france with some convenient choice I coordinate system O. The path of the particles mation in space is the particles trajectory.

plysice) loss are invarial (i.e., the same) under  
Change I institut frames. Under such a  
Godileen transformed 
$$\vec{V} = const$$
,  
 $\vec{V} = \vec{V} \cdot \vec{V} \cdot \vec{V} \cdot (t - t_0)$   
 $\vec{v} = \vec{v} \cdot \vec{V}$   
 $\vec{v} = \vec{v} + \vec{V}$   
 $\vec{v} = \vec{v} + \vec{V}$ 

Upon address of a external force 
$$\vec{F}$$
,  
a body experiences on acceleration of  
the direction of  $\vec{F}$ ,  
 $\vec{a} = \pm \vec{F}$ ,  
 $\vec{a} = \vec{v}$   
 $\vec{a} = \vec{v}$ 



time. This leads to a differential equation for the particles position as a function of time,  $\vec{r} = \prod_{m} \vec{F}(\vec{r}, \vec{r}, t)$ 

To solve this, we require and any the force Findia,  
but also some with conditions to fully specify  
the trajedary. This is usually the position  
and velocity I some reference time,  
$$\vec{r}_0 = \vec{r} |t_0|$$
 and  $\vec{v}_0 = \vec{v} (t_0)$ 

Newton's Phincipal of Determinacy The initial state of a mechanical system, i.e., the totality of possitions & velocities of its points at some moment in time, uniquely determines the all of its motion.



that is, the nation is completely determine via the integrals & nation condicioned via To & vo,

$$\vec{v}_{(4)} - \vec{v}_{o} = \int_{\vec{v}_{o}}^{\vec{v}} d\vec{v}' = \int_{m}^{t} \int_{t}^{t} dt' \vec{F}(\vec{r}_{(4')}, \vec{r}_{(4')}, t')$$

$$\vec{r}_{(4)} - \vec{r}_{o} = \int_{\vec{r}_{o}}^{\vec{r}} d\vec{r}' = \int_{t}^{t} dt' \vec{v}_{(4')}$$

NШ





Conservation Lows



We can construct a  $1^{\frac{14}{2}}$  integral it matrix by  $\vec{p}_{+} - \vec{p}_{i} = \int_{P_{i}}^{P_{p}} d\vec{p} = \int_{t_{i}}^{t_{f}} \vec{F}$   $\vec{F}_{i}$ If there are no external forces a the system, there more time is conserved!  $\vec{p}_{+} = \vec{p}_{i}$ 

The role of change in time  

$$\vec{I} = \vec{r} \times \vec{p} + \vec{r} \times \vec{p}$$
  
 $\vec{I} = \vec{r} \times \vec{p} + \vec{r} \times \vec{p}$   
 $\vec{p} = h\vec{r} \Rightarrow \vec{r} \times \vec{r} = 0$   
 $= \vec{r} \times \vec{F}$   
Define the targue  $\vec{T} = \vec{r} \times \vec{F}$ .  
Thus, we have  $\vec{I} = \vec{\Gamma}$   
Allows that it is other examinant to extra

Milie 
$$\vec{p}$$
,  $\vec{I}$  depends on the choice  $\vec{f}$  coordinde  
syder. ( $\vec{p}$ ,  $\vec{r} \rightarrow \vec{r}_0 + \vec{r} = \vec{r}'$   
 $\Rightarrow \vec{I} \rightarrow \vec{L}' = \vec{r}_0 \times \vec{p} + \vec{r} \times \vec{p} \neq \vec{l}$ ,

Every  
Cansider a force 
$$\vec{F}_{c\vec{r}}$$
, adding an a  
publick. The tworks performed on the publicle  
as it moves that  $\vec{r}_A + \vec{r}_B$  as  
 $W_{A \rightarrow B} = \int_{\vec{r}_A}^{\vec{r}_B} d\vec{r} \cdot \vec{F}_{c\vec{r}_1}$   
Note that were is in grown dependent on the  
path the product takes.  
Fin NII,  $\vec{F} = m \frac{d\vec{v}}{dt}$   
BO,  $d\vec{r} \cdot d\vec{v} = \vec{v} \cdot d\vec{v} = \frac{1}{2}d(v^2)$   
 $\Rightarrow \int_{\vec{r}_A}^{\vec{r}_B} d\vec{r} \cdot \vec{F}_{c\vec{r}_1} = \frac{1}{2}m \int_{V_A}^{V_B} d(v^2) = \frac{1}{2}m V_B^2 - \frac{1}{2}m V_B^2$   
Deliving the labor energy  $T = \frac{1}{2}m V^2$ , we  
arrive it where  $\vec{r}_B = W_{A \rightarrow B}$ 

If the force is consurdive, 
$$\nabla x \vec{F} = \vec{\sigma}$$
,  
then we can write  $\vec{F} = -\vec{\nabla} U$   
where  $U(\vec{r})$  is the poled. Jerogy.  
Gravity is an example of a consurdive force,

$$U_{grav} = -G \frac{m_i m_2}{r}$$

$$\vec{F}_{fm} = -\vec{\nabla} U_{fm}$$
$$= -G n_{i} m_{i} \hat{r}$$

We can dedue the total medianical energy E  $E = T + \sum_{j=1}^{n} O_j$ D = PEs from consording force,

So this  $\Delta E = W_{NC}$ Work from non-consume forces-If  $\vec{F}_{NC} = \vec{\sigma}$ , then  $\Delta E = cond$  energy is consumed!

System I pentides All I there notices extend to sydens I particles. We gurally brede up more of a syster & portides to the core-of-rass motion and the relative notion. Consider N particles with masses m; , j=1,..., N, and positions T; , j=1,..., N, as defined by an india system O. The color-of-mass (cm) is defined as  $\vec{R} = \pm \vec{\Sigma} m_{j}\vec{r};$ where M= Z m; is the total mass of the system. LD's apply NI to particle i,  $\vec{p}_{j} = \vec{F}_{j}^{(est)} + \sum_{k \neq j} \vec{F}_{kj}$ force, due to other periles Some entron force à le syster.

LI us examine the motion of the CM,  $M\vec{R} = \sum_{j} m_{j}\vec{r}_{j} = \sum_{j} \vec{p}_{j} = \sum_{j} \vec{F}_{j}^{(ent)} + \sum_{k \neq j} \vec{F}_{kj}$ Note that the last too can be written as  $\sum_{i} \sum_{j=1}^{n} \overline{F}_{ij} = \int_{i} \sum_{j=1}^{n} \left( \overline{F}_{ij} + \overline{F}_{jk} \right)$ whore we have defined Fij= it to reflect the absone I self forces. By NII,  $\overline{F}_{ij} = -\overline{F}_{jk}$ , s. ilaically  $\Sigma \Sigma \overline{F}_{ij} = \overline{O}$ Auctor, the che has a EOM MR = Z F (ent) = F (ent) The CM mans as a effective particle & mass M. The total romentum of the system is  $\vec{P} = \sum_{j=1}^{n} \vec{p}_{j} = d \sum_{j=1}^{n} m_{j} \vec{r}_{j} = M \vec{R} = M \vec{V}$ 

velocity & CM

The myclus moreover if the system is  

$$\begin{aligned}
\vec{L} = \sum_{j} \vec{r}_{j} \times \vec{p}_{j} \\
Table the time derivative, \\
\vec{L} = \sum_{j} \vec{r}_{j} \times \vec{p}_{j} = \sum_{j} \vec{r}_{j} \times (\vec{F}_{j}^{(eff)} + \sum_{h \neq j} \vec{F}_{hj}) \\
Now, the last tern
\\
\sum_{j,L} \vec{r}_{j} \times \vec{F}_{hj} = \frac{1}{2} \sum_{j,L} (\vec{r}_{j} - \vec{r}_{h}) \times \vec{F}_{hj}
\end{aligned}$$

When the force 
$$\vec{F}_{ij} \parallel \vec{r}_{j} - \vec{r}_{i}$$
, this is zero. Erapher  
when this is a startine is the trajedice force, when  
when needs to include EM field moresta.  
So,  
 $\vec{L} = \sum_{j} \vec{r}_{j} \times \vec{F}_{j}^{(eff)} = \nabla^{(eff)}$ .

Multiparticle Energy laws follow shill by, es;  

$$O = \frac{1}{2} \sum_{i,k} O_{kj} + \sum_{j} O_{j}^{\text{left}}.$$

Lagrangian Mechanics

Conservation leurs play central roles is physics. The Newtonian formation I mechanics often clouds the nature I conservation laws. Noether showed that conservation laws are a direct consequence I symmetries I the mechanical system under the edia I some transformation.



An attendive formulation of mechanics, called Lagrangian Hechanics, allows one to study the symptres of mechanical systems. It offers a equivalent formulation to Newton's laws. One defines the Lagrangian ZZ = T - U

The Lygrangia encodes all the dynamical information I a mechanical system. To access the equations I motion, we condrid the ation S  $S = \int_{t}^{t} Jt L$ 

By verying the coline SS=0, we can guirde the <u>Euler-Lyrange equilians</u>, which are equivalent to STI. Current a system with generalized coordinates  $E_{1}, ..., E_{n}$ , which can be position, ugles, de. By minimizing the coline, which Aledricay produces the optimal trijeday for the system, we arrive d

 $\int_{\mathcal{A}} \frac{\partial \mathcal{L}}{\partial \dot{q}_{1}} = \frac{\partial \mathcal{L}}{\partial \dot{q}_{1}}$ , ぇ=1,...,い

For example, consider the mation of a Siyse particle de some patertial well U(r). The KE I the purde is  $T = \frac{1}{2}m\dot{r}^{2} = \frac{1}{2}m\dot{z}^{2}\dot{r}^{2}$ So, h= T-U  $= \frac{1}{2}m\dot{\vec{r}}^2 - U(\vec{r})$ The Euler-Lagrange equidious yield  $\frac{\partial L}{\partial \dot{r}_{1}} = \frac{\partial}{\partial \dot{r}_{1}} \left( \frac{1}{2} n \sum_{k=1}^{2} \dot{r}_{k}^{2} \right)$ = n Z ru Suj = mrj und d 22 = nr; Hari Frugly,  $\frac{\partial L}{\partial r_i} = -\frac{2}{\partial r_j} U(\vec{r})$ Pucull  $F = -\nabla U(r)$ , so  $F_j = -2U$ hover, ⇒ mr= F